

29-Jul-2025

Term 3/Week 2

Latitude

- The main motivation for trigonometry was the astronomical observations, however, it also helped advancements in navigation.
- After recognising that the earth is spherical, it was found useful to use "angles" instead of "linear lengths" to locate the points on earth and consequently for navigation.

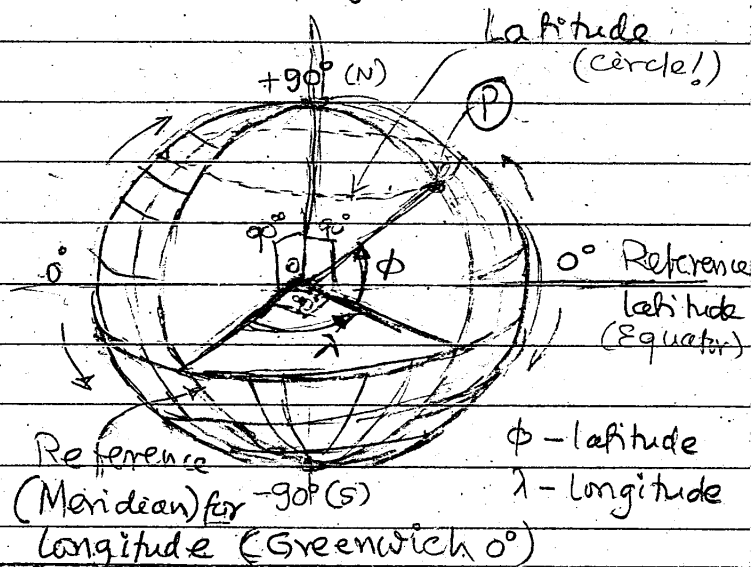
-3-

- The location of point 'P' on the sphere (earth) can be specified by angles ϕ & λ , ϕ - latitude; λ - longitude.
 - Where, ϕ is the "latitude" angle using equatorial surface as the reference. Note that angle ' ϕ ' specifies a "circle" around the earth.
- $\phi = 0^\circ \Rightarrow$ Equator
 $\phi = +90^\circ \Rightarrow$ North pole ($90^\circ N$)
 $\phi = -90^\circ \Rightarrow$ South pole ($90^\circ S$)
 $\phi = +23.5^\circ \Rightarrow$ Tropic of Cancer ($23.5^\circ N$)
 $\phi = -23.5^\circ \Rightarrow$ Tropic of Capricorn ($23.5^\circ S$)
 $\phi = (90 - 23.5) = +66.5^\circ \Rightarrow$ Arctic circle ($66.5^\circ N$)
 $\phi = (-90 - (-23.5)) = -66.5^\circ \Rightarrow$ Antarctic circle ($66.5^\circ S$)

Earth's tilt

-2-

- Latitudes & Longitudes are imaginary lines on earth's surface, which enable location and navigation.
- The best way of illustrating latitudes & longitudes is shown below.



-4-

- (Note: We will discuss Longitude details in a separate lecture).
- How to determine the latitude of a given location?
 - This can be done using a sextant and a reference star.
 - Pole star in the Northern hemisphere (easy!)
 - Southern star and southern pointer stars are used to locate south pole (Difficult)
 - Latitude can also be determined by Sun's position at Noon. But, the earth's declination angle needs to be considered.
 - Earth declination angles for Northern hemisphere seasons are:
 - Jun 22 $\Rightarrow 23.5^\circ \rightarrow$ Summer Solstice
 - Mar 21 & Sep 23 $\Rightarrow 0^\circ \rightarrow$ Equinox
 - Dec 22 $\Rightarrow -23.5^\circ \rightarrow$ Winter Solstice

Pole Star

Latitude using Pole star

- This can only be used in Northern Hemisphere.

- Angle of elevation of pole star w.r.t. (North) horizon

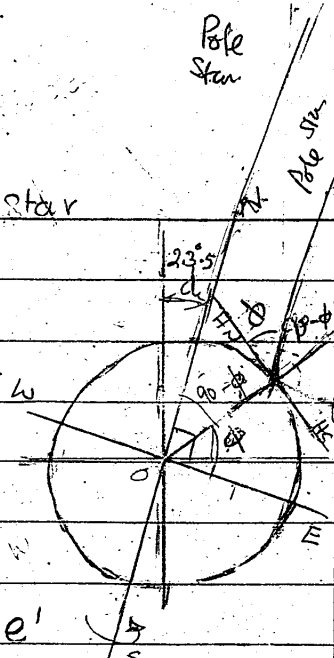
- let it be called 'e'

- Latitude of point "P"

$$\phi = e$$

- Note that earth's declension or "tilt" (23.5°) does not affect the measurement, since 'Pole star' is directly above the North Pole.

- There is no such ^{bright} star above the South Pole. ($\approx 89.35^\circ$)



- Hence, The above method is not popular in Southern hemisphere.

- A more versatile method is to measure the elevation of the Sun at Noon.

- This method is applicable for both Northern and Southern hemisphere.

- However, we need to consider Varying declension w.r.t. Sun, depending on the season.

- The declension value can be established, based on calendar days using Equinox day (say Mar 21) as our reference.

Latitude using Sun's position

- Let us consider point 'P' in the Northern hemisphere.

- The measurement is made at Noon. (Mid-day)

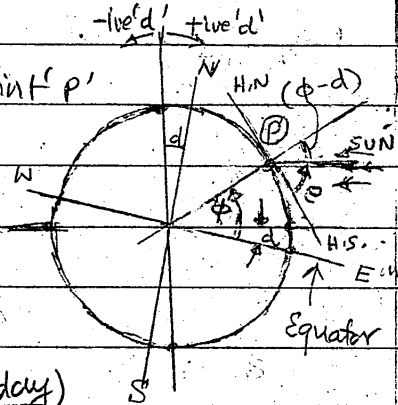
- Using Southern horizon as reference, the angle of elevation of Sun is 'e' w.r.t. (South) Horizon

- Referring to the Figure $(\phi - d) + e = 90^\circ$

- Latitude $\phi = (90^\circ + d - e)$

- Notes: (1) Declension angle (d) is positive in clockwise direction

- (2) Sun elevation angle (e) is w.r.t. Horizon-South (H-S)



- The declension angle 'd' in degrees can be calculated using the formula

$$d = 23.5 \sin \left(\frac{360}{365} \times n \right)$$

where 'n' is the number of days after Mar 21 (Northern Spring Equinox declension = 0°)

- For March 21, $n = 0$; hence declension 'd' = $0^\circ \Rightarrow$ Equinox day (Equator directly facing sun)

Alternatively,

$$d = 23.5 \sin \left(\frac{360 \times (285 + N)}{365} \right) \text{ deg}$$

where, 'N' is "Day of the Year" ($N = 1 \text{ to } 365$)

Note: Mar 21 is 80th day of the year ($\frac{285 + 80}{365}$)